

LECTURE 14: THE MEASUREMENT OF RSD

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The RSD is much more difficult to measure accurately, simply because it is harder to model the velocity power spectrum properly, even on linear scales. The commonly used model, or template, of the RSD is called the TNS model [4].

$$(1) \quad P_g(k, \mu, z) = D_{\text{FOG}}(k, \mu_2, z) [P_{g,\delta\delta}(k, z) + 2f(z)\mu^P P_{g,\delta\theta}(k, z) + f^2(z)\mu^4 P_{\theta\theta}(k, z) + A(k, \mu, z) + B(k, \mu, z)]$$

where,

$$(2) \quad \begin{aligned} P_{g,\delta\delta}(k, z) = & b_1^2(z)P_{\delta\delta}(k, z) + 2b_1(z)b_2(z)P_{b2,\delta}(k, z) \\ & - \frac{8}{7}b_1(z)(b_1(z) - 1)P_{bs2,\delta}(k, z) \\ & + \frac{64}{315}b_1(z)(b_1(z) - 1)\sigma_3^2(k, z)P_m^L(k, z) \\ & + b_2^2(z)P_{b22}(k, z) - \frac{8}{7}[b_1(z) - 1]b_2(z)P_{b2s2}(k, z) \\ & + \frac{16}{49}[b_1(z) - 1]^2P_{bs2}(k, z) \end{aligned}$$

$$(3) \quad \begin{aligned} P_{g,\delta\theta}(k, z) = & b_1(z)P_{\delta\theta}(k, z) + b_2(z)P_{b2,\theta}(k, z) \\ & - \frac{4}{7}[b_1(z) - 1]P_{bs2,\theta}(k, z) \\ & + \frac{32}{315}[b_1(z) - 1]\sigma_3^2(k, z)P_m^L(k, z) \end{aligned}$$

$$(4) \quad \begin{aligned} P_{g,\theta\theta}(k, z) &= P_{\theta\theta}(k, z) \\ D_{\text{FOG}}(k, \mu, z) &= \left\{1 + [k\mu\sigma_v(z)]^2/2\right\}^{-2} \\ A(k, \mu, z) &= b_1^3(z) \sum_{m,n=1}^3 \mu^{2m} [f(z)/b_1(z)]^n A_{mn}(k, z) \\ B(k, \mu, z) &= b_1^4(z) \sum_{m=1}^4 \sum_{a,b=1}^2 \mu^{2m} [-f(z)/b_1(z)]^{a+b} B_{ab}^m(k, z) \end{aligned}$$

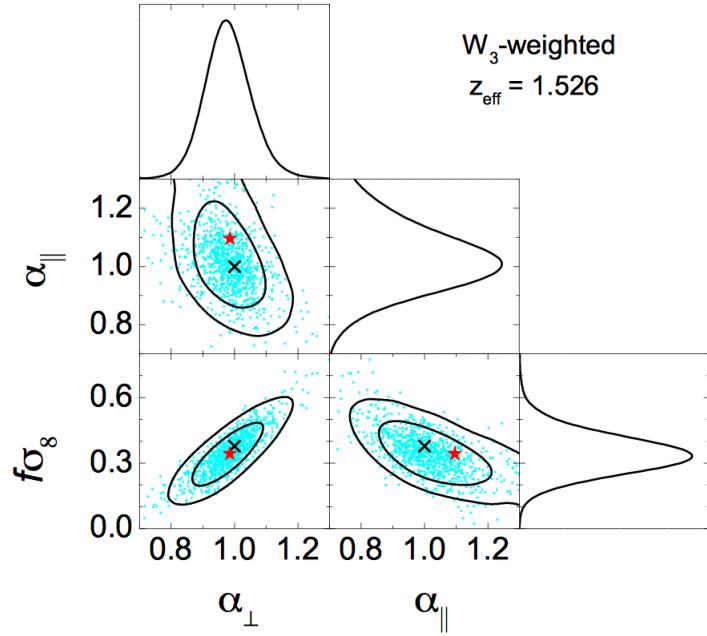


FIGURE 1. The BAO and RSD fit in k -space using BOSS DR14 quasar mock data [5].

$$(5) \quad \begin{aligned} b_{s2} &= -\frac{4}{7}(b_1 - 1) \\ b_{3nl} &= \frac{32}{315}(b_1 - 1) \end{aligned}$$

There are two approaches for the full-shape analysis: A) a fixed-template analysis and B) a full-modeling analysis.

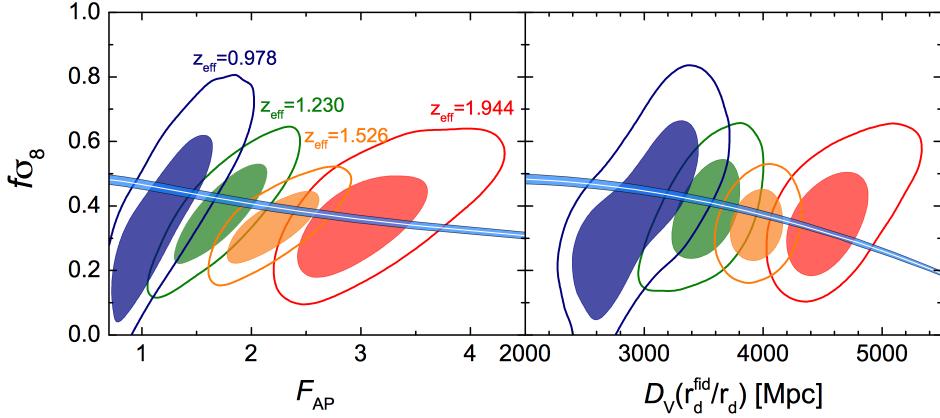


FIGURE 2. The BAO and RSD fit in k -space using the actual BOSS DR14 quasar data [5].

1. A FIXED-TEMPLATE ANALYSIS

A fixed-template analysis means that the ‘model’ used for the full-shape analysis has a fixed shape, generated using a fiducial cosmology. This is why it is called a ‘template’ instead of ‘model’. Traditional measurements for $f\sigma_8$ follow this approach, including the works in [5, 6, 7].

Recently, a new approach called ‘ShapeFit’ [8] was developed to extend the applicability of the fixed-template analysis, which can be regarded as an improved version of the fixed-template analysis (but it is still a fixed-template analysis). The new DESI analysis took this approach [9].

2. A FULL-MODELING ANALYSIS

A full-modeling analysis means that the shape of the model varies with cosmological parameters, which is more robust but computationally more expensive. In this approach, the theory model of EFTofLSS [10] is widely used. Recent implications include [9, 11, 12, 13].

Recently, a new method for the RSD measurement is developed by cross-correlating the pre- and post-reconstructed density fields [14]. This is a very efficient way to extract high-order statistics from 2-point statistics.

REFERENCES

- [1] M. J. Wilson, J. A. Peacock, A. N. Taylor and S. de la Torre, “Rapid modelling of the redshift-space power spectrum multipoles for a masked density field,” *Mon. Not. Roy. Astron. Soc.* **464**, no. 3, 3121 (2017) doi:10.1093/mnras/stw2576 [arXiv:1511.07799 [astro-ph.CO]].
- [2] G. B. Zhao *et al.* [BOSS Collaboration], “The clustering of galaxies in the completed SDSS-III Baryon Oscillation Spectroscopic Survey: tomographic BAO analysis of DR12 combined sample in Fourier space,” *Mon. Not. Roy. Astron. Soc.* **466**, no. 1, 762 (2017) doi:10.1093/mnras/stw3199 [arXiv:1607.03153 [astro-ph.CO]].
- [3] Y. Wang *et al.* [BOSS Collaboration], “The clustering of galaxies in the completed SDSS-III Baryon Oscillation Spectroscopic Survey: tomographic BAO analysis of DR12 combined sample in configuration space,” *Mon. Not. Roy. Astron. Soc.* **469**, no. 3, 3762 (2017) doi:10.1093/mnras/stx1090 [arXiv:1607.03154 [astro-ph.CO]].
- [4] A. Taruya, T. Nishimichi and S. Saito, “Baryon Acoustic Oscillations in 2D: Modeling Redshift-space Power Spectrum from Perturbation Theory,” *Phys. Rev. D* **82**, 063522 (2010) doi:10.1103/PhysRevD.82.063522 [arXiv:1006.0699 [astro-ph.CO]].
- [5] G. B. Zhao *et al.*, “The clustering of the SDSS-IV extended Baryon Oscillation Spectroscopic Survey DR14 quasar sample: a tomographic measurement of cosmic structure growth and expansion rate based on optimal redshift weights,” *Mon. Not. Roy. Astron. Soc.* **482**, no. 3, 3497 (2019) doi:10.1093/mnras/sty2845 [arXiv:1801.03043 [astro-ph.CO]].
- [6] F. Beutler *et al.* [BOSS], “The clustering of galaxies in the completed SDSS-III Baryon Oscillation Spectroscopic Survey: Anisotropic galaxy clustering in Fourier-space,” *Mon. Not. Roy. Astron. Soc.* **466**, no. 2, 2242-2260 (2017) doi:10.1093/mnras/stw3298 [arXiv:1607.03150 [astro-ph.CO]].
- [7] G. B. Zhao *et al.* [eBOSS], “The completed SDSS-IV extended Baryon Oscillation Spectroscopic Survey: a multitracer analysis in Fourier space for measuring the cosmic structure growth and expansion rate,” *Mon. Not. Roy. Astron. Soc.* **504**, no. 1, 33-52 (2021) doi:10.1093/mnras/stab849 [arXiv:2007.09011 [astro-ph.CO]].
- [8] S. Brieden, H. Gil-Marín and L. Verde, “ShapeFit: extracting the power spectrum shape information in galaxy surveys beyond BAO and RSD,” *JCAP* **12**, no. 12, 054 (2021) doi:10.1088/1475-7516/2021/12/054 [arXiv:2106.07641 [astro-ph.CO]].
- [9] A. G. Adame *et al.* [DESI], “DESI 2024 V: Full-Shape Galaxy Clustering from Galaxies and Quasars,” [arXiv:2411.12021 [astro-ph.CO]].
- [10] J. J. M. Carrasco, M. P. Hertzberg and L. Senatore, “The Effective Field Theory of Cosmological Large Scale Structures,” *JHEP* **09**, 082 (2012) doi:10.1007/JHEP09(2012)082 [arXiv:1206.2926 [astro-ph.CO]].
- [11] G. D’Amico, L. Senatore and P. Zhang, “Limits on w CDM from the EFTofLSS with the PyBird code,” *JCAP* **01**, 006 (2021) doi:10.1088/1475-7516/2021/01/006 [arXiv:2003.07956 [astro-ph.CO]].
- [12] A. G. Adame *et al.* [DESI], “DESI 2024 VII: Cosmological Constraints from the Full-Shape Modeling of Clustering Measurements,” [arXiv:2411.12022 [astro-ph.CO]].
- [13] M. Ishak, J. Pan, R. Calderon, K. Lodha, G. Valogiannis, A. Aviles, G. Niz, L. Yi, C. Zheng and C. Garcia-Quintero, *et al.* “Modified Gravity Constraints from the Full Shape Modeling of Clustering Measurements from DESI 2024,” [arXiv:2411.12026 [astro-ph.CO]].
- [14] Y. Wang, G. B. Zhao, K. Koyama, W. J. Percival, R. Takahashi, C. Hikage, H. Gil-Marín, C. Hahn, R. Zhao and W. Zhang, *et al.* “Extracting high-order cosmological information in galaxy surveys with power spectra,” *Commun. Phys.* **7**, no. 1, 130 (2024) doi:10.1038/s42005-024-01624-7 [arXiv:2202.05248 [astro-ph.CO]].